

RESEARCH ARTICLE | MAY 10 2024

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Physics of Fluids 36, 056111 (2024)

<https://doi.org/10.1063/5.0199992>



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Cite as: Phys. Fluids **36**, 056111 (2024); doi: [10.1063/5.0199992](https://doi.org/10.1063/5.0199992)

Submitted: 24 January 2024 · Accepted: 23 April 2024 ·

Published Online: 10 May 2024



View Online



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J. N. Fréreau,¹ B. Tomasetti,¹ A. Benidar,² G. Liger-Belair,³ A. B. Himakar,⁴ M. P. Lakhan,⁴ R. P. Srikar,⁴ S. V. Nikhilesh,⁴ J. Copalan,⁴ S. M. V. Rao,⁴ R. Georges,^{2,a)} and C. Lauzin¹

AFFILIATIONS

¹Institute of Condensed Matter and Nanosciences, Université Catholique de Louvain, 1348 Louvain-la-Neuve, Belgium

²Institut de Physique de Rennes, UMR CNRS 6251, Université de Rennes, Campus de Beaulieu, 35042 Rennes Cedex, France

³Groupe de Spectrométrie Moléculaire et Atmosphérique, UMR CNRS 7331, Université de Reims Champagne-Ardenne, 51687 Reims, France

⁴Laboratory for Hypersonic and Shock wave Research, Indian Institute of Science, Bengaluru, Karnataka 560001, India

^{a)}Author to whom correspondence should be addressed: robert.georges@univ-rennes.fr

ABSTRACT

Cork popping represents a rich and complex fluid dynamics process, involving up to three phases (liquid, gas, and solid), three main chemical compounds (ethanol, water, CO₂), and a moving cork gradually opening the bottle and blocking the fast progression of the expanding gas. In this work, we used high-speed, high-sensitivity schlieren imaging setups to provide a fresh perspective on this dynamical phenomenon. Our experimental results are systematically compared and interpreted on the basis of new computational fluid dynamics simulations. Our combined experimental and numerical works confirm the establishment of two supersonic expansions during cork popping from a champagne bottle.

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I. INTRODUCTION

From a physico-chemical point of view, champagne and other sparkling wines can be viewed as multicomponent hydroalcoholic solutions. They present a surface tension $\gamma \approx 50 \text{ mN m}^{-1}$ (i.e., reduced by about a third compared to that of pure water, mainly due to $\approx 12\%$ – 13% ethanol by volume), a viscosity close to 1.5 mPa s ($\approx 50\%$ more than that of pure water mainly because of ethanol), and a density relative to pure water close to unity.¹ Moreover, champagne and other sparkling wines are saturated with dissolved carbon dioxide (CO₂), whether during a second in-bottle alcoholic fermentation process for premium sparkling wines like champagne, or through simple exogenous gas-phase CO₂ injection for some cheaper sparkling wines.² For premium sparkling wines, the CO₂ produced by the yeasts during this second in-bottle fermentation cannot escape and progressively dissolves into the wine.

In a sealed bottle of sparkling wine, gas-phase CO₂ and dissolved CO₂ undergo thermodynamic equilibrium according to the so-called Henry's law. The concentration c_L of dissolved CO₂ in the wine is indeed proportional to the partial pressure of gas-phase CO₂,

according to the formula $c_L = k_H P_{\text{CO}_2}$, with k_H being the strongly temperature-dependent Henry's constant of gas-phase CO₂ in wine³ and P_{CO_2} being the partial pressure of gas-phase CO₂ in the sealed bottle. In addition, the temperature dependence of Henry's constant leads to a strong temperature dependence of the pressure prevailing in the bottle hermetically sealed with a cork as it was modeled by Liger-Belair *et al.*⁴ It goes from roughly 5 bars at 8 °C to more than 10 bars at 30 °C.^{4,5} During the cork popping process, concomitantly with the cork stopper being expelled from the bottleneck under the action of CO₂ pressure, the underexpanded gas of the champagne bottle (mainly composed of CO₂ with traces of water vapor and ethanol) expands out of the bottleneck through the ambient air and experiences adiabatic cooling.^{4,5}

While free underexpanded jets are very well documented, for example see Franquet *et al.*,⁶ expansions that are obstructed are still challenging to rationalize.^{7–9} In this respect, the cork popping of a champagne bottle is particularly interesting and challenging, as the position of the obstacle changes dynamically as the gas expands. Moreover, it involves up to three phases (gas, liquid, and solid particles

formed during severe adiabatic cooling) as well as a mixture of gases with different physico-chemical characteristics. This cork popping dynamics was recently studied using high-speed imaging.⁵ Films were recorded for bottles stored at 20 and 30 °C. These recordings allowed the observation of a planar shock wave (Mach disk) and of partial nucleation of H₂O and/or CO₂. The amplitude of this nucleation was different for the two temperatures and was deduced from the change of color following the major light scattering mechanism at play, i.e., Rayleigh or Mie scattering. This work initiated a numerical effort to better understand the details of the fluid dynamics involved.¹⁰ A computational fluid dynamics (CFD) method was then used to simulate the expansion of 7.5 bars of pure CO₂ obstructed by a rigid cork moving at a constant speed. The authors, most of them involved in this work, computed dynamically the Mach number, temperature, pressure, and density of the expanding gas with a time resolution of 1 μs. The simulation predicted the apparition of a complex set of radial shock waves at early time of the expansion. The experimental observation of a planar shock wave going first toward the cork and later, when the bottle pressure decreased, toward the bottle was also reproduced by the modeling. Diamond-shaped cells expanding radially from the compressed gas impacting the cork were predicted but not observed experimentally. In this work, we use high-speed schlieren imaging combined with new CFD simulations to provide a new look into this interesting phenomenon.

The article is organized as follows: The physical characteristics of typical champagne bottles are presented in Sec. II A. The improvements of the CFD simulations compared to the work of Benidar *et al.*¹⁰ are provided in Sec. II B. The experimental setups and procedures are presented in Sec. II C. The results of the CFD simulations are described in Sec. III A to present the experimental results obtained through high-speed schlieren imaging in Sec. III B. Finally, conclusions and perspectives are discussed in the last section of this paper.

II. MATERIALS AND METHODS

A. Presentation of the physical parameters of a typical champagne bottle

Experimentally, a series of 75 cL bottles obtained from different suppliers were used. Typically, champagne bottles present 12%–13% ethanol by volume concentration. To trigger the second in-bottle alcoholic fermentation needed to force the dissolution of gaseous CO₂ in these bottles (the so-called *prise de mousse*), yeasts along with 20–24 g L⁻¹ of sugar were classically added in the bottles hermetically sealed. The key metabolic process behind the production of gas-phase CO₂ in the sealed bottles is alcoholic fermentation. Actually, 20–24 g L⁻¹ of sugar added in the blend leads to the production of 7.3–8.8 g of CO₂ (i.e., ≈ 0.17–0.2 mol) in every bottle of champagne.⁴ After the *prise de mousse* was achieved, bottles are aged in a cool cellar for at least 15 months. Bottles are then disgorged, which consists in opening the bottles to remove the dead yeast cells and recorking them with traditional natural cork stopper. During this step, the thermodynamic equilibrium of CO₂ is broken, but the dissolved and gas-phase CO₂ quickly recover Henry's equilibrium in the recorked bottle where a volume V_G of gas phase (the headspace) still cohabits with a volume V_L of champagne (i.e., the liquid phase). It is worth noting that, into every bottle of a batch, the volume V_G of the headspace under the cork is 25–35 ml, whereas the volume V_L of champagne is 75 cL. Recently, the pressure-temperature relationship in the corked bottles (after the

disgorging step was achieved) was theoretically determined.⁴ It obeys the following relationship, with every parameter being expressed in the international system of units (SI):

$$p^{\text{CO}_2} = \frac{nk_H(RT)^2 V_L}{(V_G + nk_H RT V_L)^2}, \quad (1)$$

with p^{CO_2} being the CO₂ pressure in the corked bottle, n being the total number of CO₂ moles produced in a sealed bottle after the *prise de mousse* (i.e., 0.17–0.2 mol in the present case), k_H being the temperature-dependent Henry's constant of gas-phase CO₂ in wine,³ R being the ideal gas constant, and T being the temperature. Thermodynamically speaking, the temperature-dependent Henry's constant of gas-phase CO₂ in wine was found to be conveniently expressed with a van't Hoff-like equation as follows:³

$$k_H(T) = k_{298K} \exp \left[-\frac{\Delta H_{\text{diss}}}{R} \left(\frac{1}{T} - \frac{1}{298} \right) \right], \quad (2)$$

with k_{298K} being the Henry's constant of CO₂ at 25 °C (≈ 1.2 g L⁻¹ bar⁻¹, i.e., 2.7×10^{-4} mol m⁻³ Pa⁻¹) and ΔH_{diss} being the dissolution enthalpy of CO₂ in wine (≈ -24 800 J mol⁻¹). For the standard commercial champagne bottles of the present study, applying Eq. (1) leads to CO₂ pressures ranging between 7.4 and 8.9 bars in the corked bottles stored at 25 °C.

B. CFD calculations

Our numerical simulations represent an extension of a previous investigation presented in Benidar *et al.*¹⁰ We recommend the readers to refer to our initial publication for comprehensive details on the basic framework and associated assumptions. Briefly, all the simulations are based on the use of the commercial solver Ansys Fluent. These simulations are conducted under the assumption of axisymmetric geometry. The standard k-epsilon turbulence model is employed with a near-wall treatment. This turbulent model was implemented in order to deal with large velocity fluctuations in the region of radial expansion and in the shear layers of axial expansion (colormaps of turbulence intensity are available in the supplementary material, Fig. S1). The k-ε Reynolds-averaged Navier–Stokes (RANS) model is preferred to more elaborate models dedicated to the dynamic interactions of fluids around the moving object (such as the unsteady RANS model), in order to balance accuracy and computational feasibility. While this approach may not capture all the transient effects, it provides a reasonable approximation of the flow behavior and allows us to obtain valuable insights within the constraints of our computational resources.

For simplicity, we assume that the gas contained in the bottle is composed exclusively of CO₂. Initial conditions are defined with a 35 ml headspace, a pressure of 7.5 bars, and a temperature of 20 °C. The external gas is air at 20 °C and atmospheric pressure. According to current and past experimental observations,^{4,5} no net variation in cork velocity was observed after ~100 μs and thus for most of the duration of supersonic expansions. Consequently, the cork velocity is fixed as a constant (18.6 m s⁻¹) in the numerical simulations. A CAD (computer-aided design) file of a real champagne bottle was used to define the bottle profile. Moreover, it is assumed that the cork geometry is not modified during the process. This assumption simplifies the simulation and is realistic for old champagne bottles. Indeed after 12 years of age, a natural cork stopper deforms less and tends to keep a

cylindrical shape whose diameter corresponds to that of the bottle neck¹¹ (more information available in the supplementary material). For a detailed study of cork deformation during uncorking and its consequences on the gas dynamics, readers should refer to Wagner *et al.*¹²

The quality of CFD simulations compared to our previous study¹⁰ was improved by enhancing the spatial and temporal resolution. This was performed by decreasing the calculation time step from 1 μ s to 50 ns and by changing our meshing strategy. To ensure precise predictions of the flow field, it is indeed imperative to provide an adequate grid density in the mixing region. This includes sufficient grid density to capture high-velocity gradient regions within the shear layer and a large number of grid elements to resolve high-pressure gradient regions near shock fronts. In our prior study, which employed a dynamic meshing technique, grid density in the mixing region proved highly sensitive to dynamic layering, smoothing, and remeshing techniques. The efforts to refine the mesh in a specific flow region often resulted in calculation divergence.

In these revised numerical simulations, we have adopted an overset meshing method^{13,14} capable of accommodating moving bodies within turbulent flows. Unlike the previous approach that involved sliding meshes, mesh morphing, and remeshing, this alternative mesh handling strategy creates separate grids that overlap instead of being complementary. Each component is treated as an independent body with its dedicated grid. The overset method simultaneously computes all grids, encompassing the component zones and background zones. Field variables are interpolated from all grids at each time step. This approach mitigates computational complexity and ensures realistic motion simulations for complex geometries. As previously mentioned, the computational domain is divided into two parts: an overset mesh comprising the moving cork and the surrounding fluid, and a background mesh used to model the static bottle and its surrounding fluid, as depicted in Fig. 1. In both domains, the meshing is variable, ranging from 0.001 mm for the finer mesh to 1 mm for the coarser mesh. Cell zone conditions are specified with a fixed mesh for the background grid and a moving mesh for the component grid, featuring a translational velocity of 18.6 m/s. Boundary conditions for walls are set as stationary for the bottle and as moving for the cork with a velocity of 18.6 m/s. Shear conditions for all walls are defined as non-slip.

C. Experimental apparatus

Two experimental setups were used in this work, one located at UCLouvain (Louvain-la-Neuve, Belgium) and the other at the Laboratory for Hypersonic and Shock wave Research (Bangalore, India). These setups will be called hereafter UCLouvain and LHSR, respectively. The general scheme of the two setups is the same and is presented in Fig. 2. This corresponds to a z-type parallel-light schlieren system.^{15–18} In UCLouvain, a halogen lamp is used in combination with a short focal lens and a diaphragm of 3 mm diameter to produce a point (white) light source (LS). This point source is placed at the focal point of a first spherical mirror (SM1) ($D_1 = 0.46$ m, $f_1 = 5$ m). The resulting collimated beam passes through the studied area and a second spherical mirror (SM2) ($D_2 = 0.70$ m, $f_2 = 3.5$ m), placed 2.6 m after the studied area, images the light beam source on its focal point where a spatial filter is placed. Light transmitted by the spatial filter is recorded by a camera (HighSpeedStar 3 G, LaVision). Camera recording parameters are set to 20 000 frames s^{-1} with a 512×128 pixel matrix (8-bit gray scale) and 50 μ s exposure time.

In LHSR, the two mirrors have focal lengths of $f_1 = f_2 = 3$ m and diameters of $D_1 = D_2 = 280$ mm. The light source is a high power LED (operated at 40 W) emitting continuous light at 532 nm. The camera used for the acquisition is a Photron FASTCAM SA4, recording 512×304 pixels images at a rate of 22 500 frames per second and 3.25 μ s exposure time. For all measurements presented in this work, spatial filtering was performed using a razor blade (RB) knife-edge in order to maximize the contrast on supersonic patterns. The spatial filtering was performed vertically, i.e., the blade is placed along the x axis, since the vertical axis y is the axis of the expansion (bottom part of light spot is cut by the RB).

In UCLouvain, a lens (L2 in Fig. 2) with a long focal length of $f = +10$ m is placed a few centimeters from the studied area and allows the quantification of the total deviation angle of the light¹⁷ as detailed in the supplementary material (Fig. S1). Using this calibration, a minimum measurable deviation was evaluated to be 0.4 arc sec (100 μ deg) at a distance of 6.1 m between the lens and the RB. Total light deviation can be used to quantify local thermodynamic variations.^{19–22} The estimation of the total deviation is additional information given by our studies but is not developed in this paper. All the figures presented in this work are from two measurement campaigns carried out at

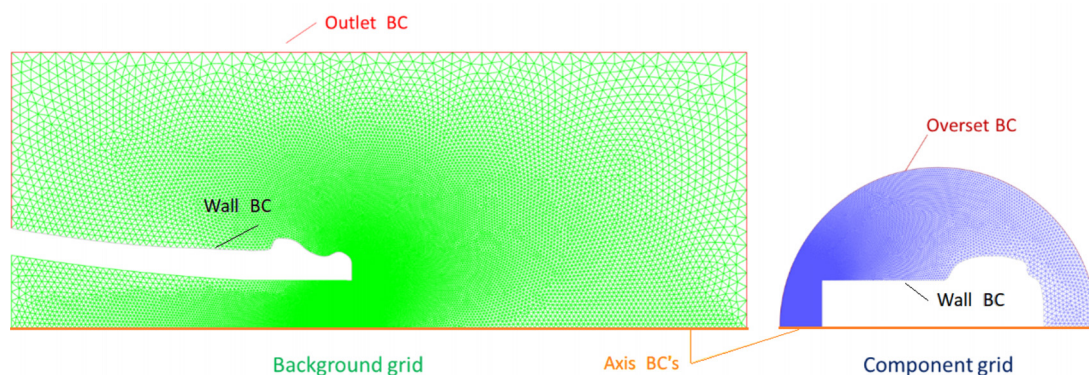


FIG. 1. Overset mesh and boundary conditions (BCs). Two distinct computational domains are defined: the background domain (length $\times$ width: 160×45 mm²) and the overset domain (radius: 30 mm).

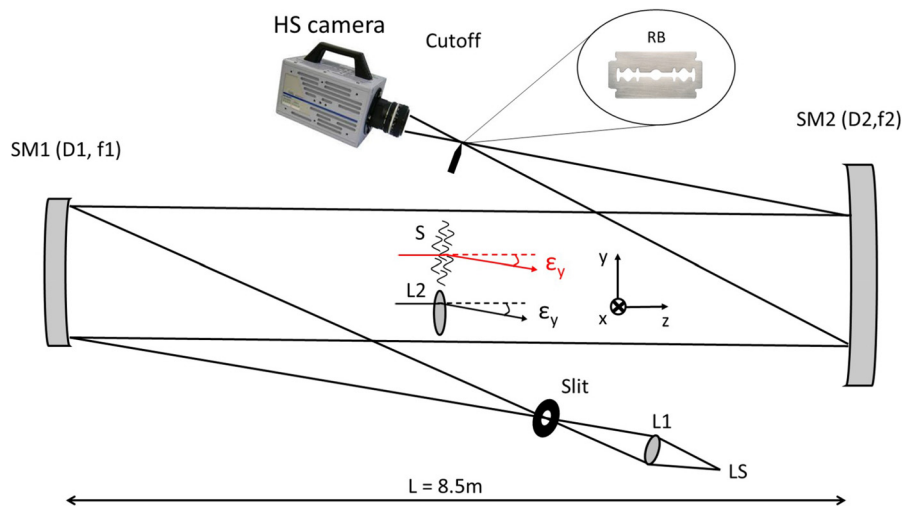


FIG. 2. Schematic of the parallel-light z-type schlieren setup. SM stands for spherical mirror, D for diameter, f for focal length, LS for light source, S for sample, and ε for light deviation.

UCLouvain in December 2022 and June 2023, with the exception of the results presented in Fig. 4 associated with measurements carried out at LHSR in November 2022. For measurements done at UCLouvain, the champagne and sparkling wine bottles were placed and maintained at a given temperature either using a bain-marie and a heating plate or using a heating tape and a PID device to reach a reference temperature.

III. RESULTS AND DISCUSSION

The results of the CFD simulation (Fig. 3) and those of the experiments carried out at LHSR and UCLouvain (Figs. 4 and 5, respectively) have multimedia content available (Multimedia view).

A. CFD simulations results

The CFD predictions of the flow density evolution are presented in Fig. 3. The full animation is also available (Multimedia view). This dynamics can be arbitrarily divided in three stages for times ranging from 0 to 1.5 ms, where $t = 0$ is defined as the frame before the appearance of exiting gases. The first stage of the popping process is illustrated in panel (a) of Fig. 3. A radial expansion through the thin annular gap is created when the cork begins to separate from the bottle. An expansion zone (rarefaction fan), typical of free underexpanded jets, forms at the edges of the bottleneck, and Mach number simulation (not shown) reveals that the gas has already acquired supersonic radial velocity.

The second stage is illustrated in panels (b) and (c) of Fig. 3. The gas expansion just above the bottleneck becomes an axial jet whereas the start of radial jet remains located at the base of the cork. Mach number simulation (not shown) reveals that the axial expansion is also supersonic. However, the cork clearly compresses the impinging gas and therefore this expansion can be defined as an impinging supersonic jet.^{8,23–27} The formation of a first Mach disk is predicted as a sudden change of density is observed when the gas flows from the axial expansion to the compression zone under the cork. In this regime, the impinging gas maintains a sufficiently high pressure ratio in the compression zone compared with atmospheric pressure: The radial jet remains supersonic, as shown by the succession of diamond shock

cells. Therefore, the gas flow from the bottle undergoes two distinct supersonic expansions: first, an axial impinging supersonic expansion at the bottle exit, followed by a compression of the gas due to presence of the cork and, second, a re-acceleration to supersonic speed during the radial free expansion. The distance of the Mach disk from the bottle increases as the standoff distance between the cork and the bottle increases and reaches a maximum that defines the ends of stage no. 2. In our previous CFD simulation,¹⁰ the impact of the cork on the position of the Mach disk was described, emphasizing the impinging nature of the axial expansion.

The third stage associated with $t \geq 800 \mu\text{s}$ is presented in panels (d)–(h) of Fig. 3. The pressure ratio between the compression zone under the cork and the bottle drops below a critical point where the extension of the axial supersonic expansion cannot be sustained. The Mach disk formed by the axial expansion appears to fall back to the bottle.^{28,29} During the collapse of this first Mach disk, a second Mach disk is created further downstream of the axial expansion (i.e., at a larger distance from the bottleneck). At a later time, this second disk appears to follow the first one by collapsing toward the bottle. The second Mach disk [Fig. 3, panel (e)] is delimiting the end of the second shock cell. It should be noted that the drop in bottle pressure leads to a continuous decrease in pressure ratio, which shortens the shock cells and moves the Mach disks toward the bottleneck.^{7,8} At the same time, the supersonic character of the radial expansion is lost as the pressure under the cork is no longer sufficient to sustain it. Finally, instabilities are developing in the boundary layers of the jet (vortices) and a compression wave propagates inside the bottle [Fig. 3, panel (h)].

The improvement of the temporal and spatial resolution of the CFD simulations allowed to provide new information. Especially the appearance of a second Mach disk during the collapse stage is a new feature directly associated with these improvements. This pattern was also numerically predicted recently in Wagner *et al.*¹² A possible limitation of this CFD model can also be mentioned. In panels (d) and (e) of Fig. 3, the simulations predict the appearance of large vortices in the radial expansion after the loss of its supersonic character. However, as we shall see in Sec. III B, these vortex-like structures have not been observed experimentally so far and may be numerical artifacts. It

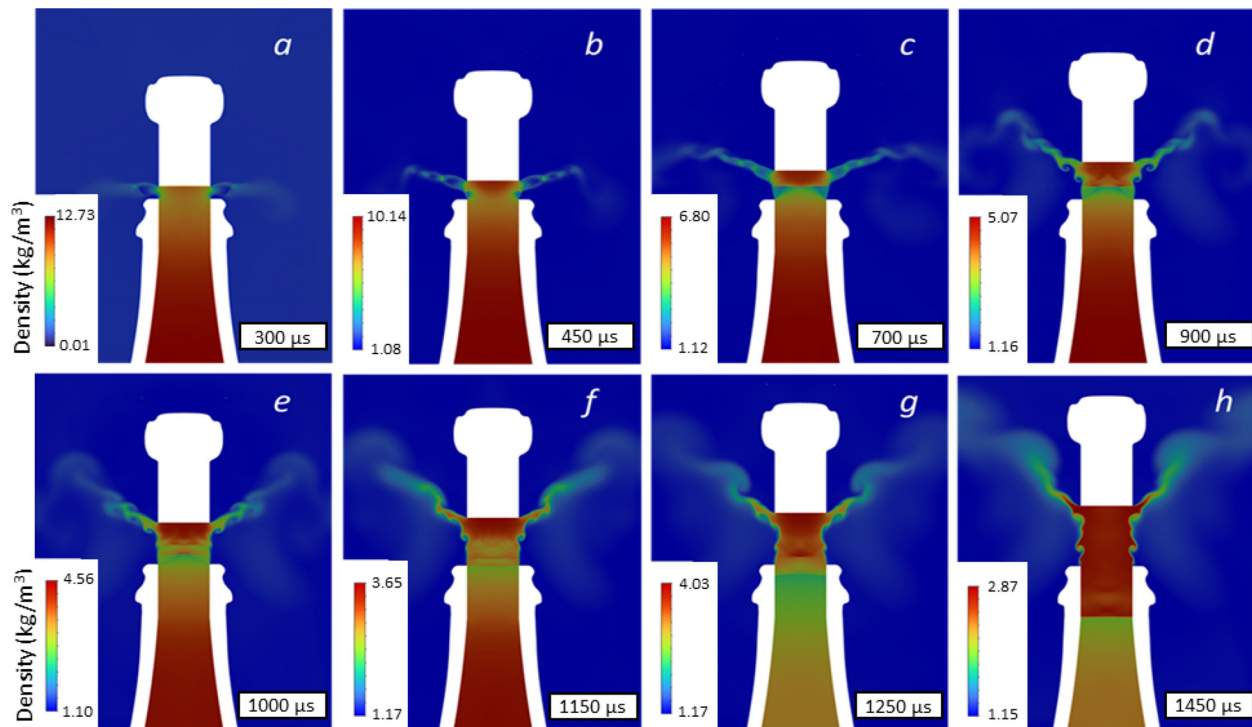


FIG. 3. CFD simulation of the 2D axisymmetric density profile for the cork popping process using the overset meshing method. The initial parameters are as follows: headspace gas volume = 35 ml, pressure = 7.5 bar, and temperature = 20 °C. The gas in the bottle is pure CO₂ and the surrounding gas is air. A variable density scale is used to enhance the contrast of the fluid dynamics features. The same figure with a fixed scale is provided in the supplementary material. Multimedia available online.

should be mentioned that the presence of vortex-like structures is only indicative in schlieren images, which is limited by line-of-sight integration effects and camera spatial resolution. Improvements in experimental spatial resolution and combination with other types of measurement, such as particle image velocimetry, would be required to observe such features.

B. High-speed schlieren imaging of a cork popping off a champagne bottle

Experimentally, our first objective was to confirm using schlieren imaging the presence of the normal shock wave (Mach disk) first observed by visible light scattering by Liger-Belair *et al.*⁵ and subsequently confirmed by numerical simulations.^{10,12} The LHSR experimental setup was used to visualize gas ejection from the head volume of a bottle of sparkling wine held at 25 °C. Figure 4 shows a chronological sequence of selected images filmed at the LHSR; the complete film is available (Multimedia view), $t=0$ is defined as (previously) the frame before the appearance of exiting gases. The film clearly reveals the formation of a Mach disk (MD1) around 711 μ s after the cork separates from the bottle (remember here that the time interval between two images is 44.4 μ s, so MD1 could in fact have appeared earlier in a 44.4 μ s time interval). This observation is nicely confirming the supersonic nature of the axial gas expansion and the relevance of the numerical simulations in Benidar *et al.*¹⁰ and in Wagner *et al.*¹² MD1 then rises steadily above the bottle, before suddenly deepening toward the

bottle at around 1289 μ s as predicted in numerical simulations. A second Mach disk (MD2) forms downstream of MD1 at 1422 μ s, slightly beyond the highest position ever reached by MD1 and then falls toward the bottle as predicted by the CFD simulation. This is the first experimental observation of the formation of this second Mach disk in the cork popping process. At 1511 μ s, MD1 disappears into the bottleneck, followed by MD2 at 1600 μ s. The graph in the last panel of Fig. 4 shows the temporal evolution of the position of MD1 and MD2, as well as the base of the cork relative to the bottleneck. The splitting of points at short times is associated with the observation of a thick disk instead of a line at early times of the expansion: Each series of points is associated with the maximum and minimum height of this disk.

A second measurement campaign was carried out at UCLouvain, during which new fluid dynamics phenomena were revealed. Specific frames of the dynamics of a champagne popping during this latter campaign are displayed in Fig. 5 (Multimedia view). The time, t , associated with each panel is indicated in the lower left side. This was recorded with a frame rate of 20 000 fps, i.e., $\Delta t = 50$ μ s time resolution. Once again, $t=0$ is defined as the frame before the appearance of exiting gases. The velocity of the expelled cork is rapidly increasing reaching a value of 17 m/s after $t = 150$ μ s, a value similar to the one obtained in previous studies.⁵ Similar to the results above, the rise and fall of the horizontal Mach disk was again observed (see 600, 750, 900 μ s in Fig. 5). However, the characteristic times associated with this process are clearly different as the maximum stretch of Mach disk 1 was occurring 1200 μ s after the first appearance of gas in Fig. 4, whereas

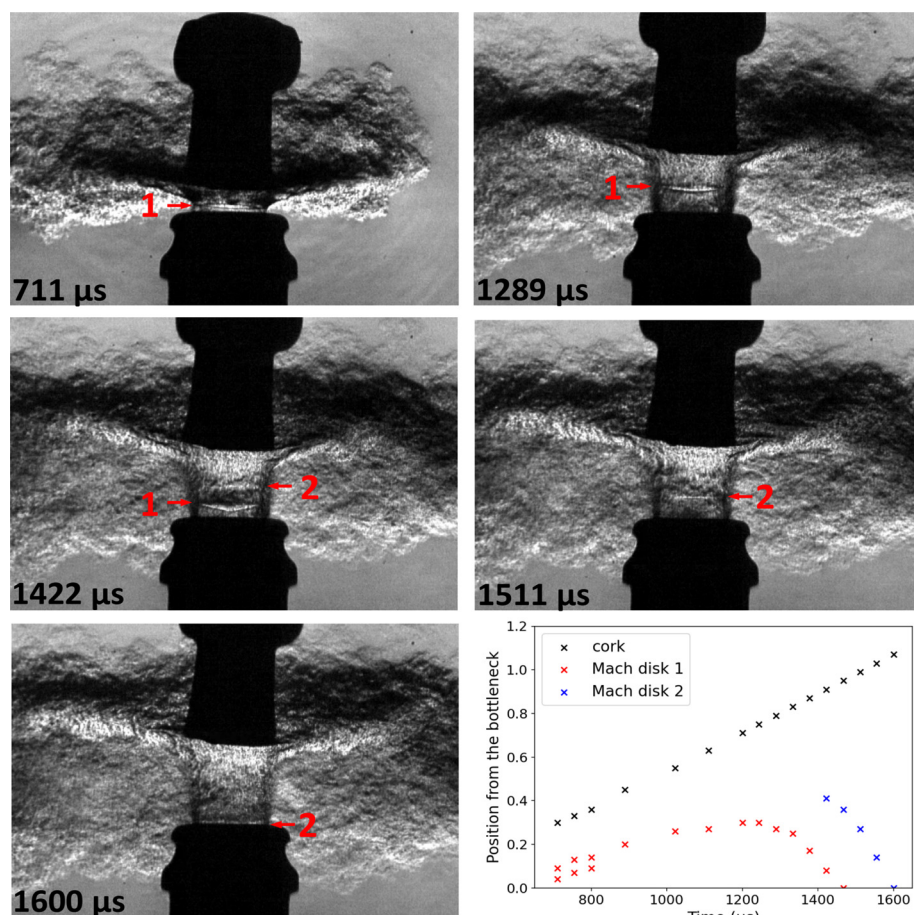


FIG. 4. Time sequence showing the cork popping process of a champagne bottle at 25 °C at LHSR in Bangalore, India. The times shown on the frames correspond to the times elapsed after the first appearance of gas. Spatial filtering was performed using a RB placed horizontally. The last panel shows the positions of the cork (black), the Mach disk 1 (red), and Mach disk 2 (blue) vs time. Positions are expressed in diameter of the bottleneck. Multimedia available online.

the global supersonic dynamics is already finished at that time in Fig. 5. This time dispersion was already observed previously⁵ and is characteristic of changes in thermodynamic conditions and, in particular, changes in temperature. Nevertheless, despite this time dispersion, the fluid dynamics features predicted by CFD calculations were observed in these films.

Fluid dynamics of the cork popping process evolves rapidly over time. At early times, e.g., at $t = 300, 450, 600,$ and $750 \mu\text{s}$, lateral diamond shock cells are clearly visible. The origin of these structures is located at the basis of the cork and follows its movement. These structures were predicted by CFD modeling^{10,12} but were so far never observed experimentally. This new experimental observation confirms

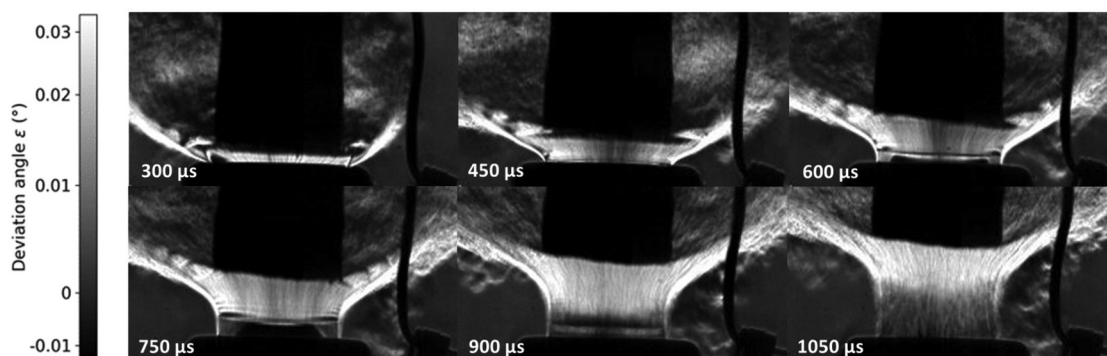


FIG. 5. Time sequence showing the cork popping process of a champagne bottle at 25 °C in UCLouvain, Belgium. The time of each frame is indicated in the lower left. Spatial filtering was performed using a razor blade placed horizontally. Multimedia available online.

the formation of a radial free supersonic jet as discussed in Sec. III A. The lateral shock cells feature remains clearly visible despite the latter formation of a Mach disk (e.g., $t = 600$ and $750 \mu\text{s}$) through which the axial flow at the bottle outlet becomes subsonic close to the cork. An interesting frame is the one at $t = 750 \mu\text{s}$. This frame is reproduced in Fig. 6 with a comparison with CFD simulation of the spatial distribution of the Mach number. The comparison is not direct but illustrative since it presents two different quantities: the Mach number (CFD simulation) for the upper panel and the density gradient along the y axis for the lower panel (schlieren image). In the simulation, the color codes clearly indicates subsonic (white to black) and supersonic (green to red) regions. Five zones are highlighted in the experimental image and discussed in detail below.

In zone 5, the above-mentioned radial expansion is clearly visible through periodic pattern of alternating dark and bright areas on both sides of the expansion. This regular pattern develops very early in the process, at $t = 150\text{--}200 \mu\text{s}$, and is characteristic in schlieren imaging of a diamond structure originating from underexpanded supersonic jets.^{30–32} Diamond-shaped cells are observed at the start of the expansion stage and last until $t = 800 \mu\text{s}$, when the cork is 1 cm above the bottle neck (see Fig. 5 at $750 \mu\text{s}$). The observation of diamond-shaped shock cells at this later stage requires the presence of a high-pressure gas just under the cork (zone 4 in Fig. 6). This is in line with the expansion scheme described above in Sec. III A where a high-pressure zone under the cork is formed. In addition, as observed in the experimental image, the radial deformation of the cork is very limited. Our numerical simulations, assuming a non-deforming cork, are able to reproduce the deflection angle of the radial expansion. Interestingly, this angle of deflection is much more pronounced in the images in Fig. 4 with a clear expansion of the cork once the bottle is fully open. This

deformation was taken into account in the simulations carried out by Wagner *et al.*¹² (see Fig. S8) with a mentioned relation between the radial expansion of the cork during the opening and the large deflection angle of the radial expansion. This is thus confirmed by our observations.

Zone 3 highlights a thin horizontal feature located respectively upstream of zone 4 and downstream of zone 1. A similar feature is observed in high-speed direct color imaging.⁵ This feature was also predicted in our simulations and in Wagner *et al.*¹² and is ascribed to the formation of a horizontal Mach disk.

The dark zone 1 shown in Fig. 6 is located upstream of the first Mach disk. Zone 1 is in sharp contrast with the surrounding area (zone 2 in Fig. 6). This is an indicator of an edge in gas density variation. In this experimental measurement, the spatial filtering (i.e., the razor blade) is cutting the bottom part of the light spot. Therefore, dark areas in our images correspond to negative gradient of the refractive index (and thus of density) along the y axis (see the supplementary material, Fig. S4). As zone 1 is appearing black in our images, zone 1 is a higher pressure area when compared to zone 2, so we assigned zone 1 to the subsonic part of the expansion and zone 2 to the supersonic zone. This pure experimental assignment is in agreement with the CFD maps shown in Fig. 6. In these CFD maps, a convex sonic line is expected and the axial extension of the supersonic area is maximized on the side of the expansion and not on the central axis. Convex shape of the subsonic area at the exit of the nozzle is commonly observed in supersonic expansions.^{33,34} Experimental observation of the features corresponding to zones 1 and 3 (Mach disk) confirms the supersonic nature of the axial expansion.

In zone 4, the compression zone shows alternating light and dark horizontal ripples just above zone 3. These ripples look blurry in Fig. 6, but they are clearly visible in the video (Multimedia view). These ripples appear over time as the main Mach disk rises. The number of oscillations increases over time until they disappear completely very shortly (1 frame or around $50 \mu\text{s}$) after the main Mach disk begins to descend toward the bottleneck. Therefore, the presence of these ripples seems to be strongly linked to that of the main Mach disk. On the CFD map dedicated to Mach number in Fig. 6, small supersonic cells are computed on the side of the compression zone. The supersonic cells are in fact ring-shaped and are probably responsible for the observation of these ripples in the experimental image. These ring-shaped supersonic cells were predicted neither in our previous CFD¹⁰ nor in reference Wagner *et al.*¹² However, our updated calculation method now allows us to predict similar structures, which could be related to reflections of compression waves after the Mach disk. A viscous instability could be another explanation, which would explain why our CFD simulations that include viscosity manage to partially reproduce these periodic structures, whereas the simulations in Wagner *et al.*¹² that neglect viscosity do not. A more detailed comparison between the experiment and the simulations about this zone 4 can be found in the supplementary material, Fig. S5.

The time-dependent positions of the cork and the first Mach disk measured in the present work are compared in Fig. 7 with those observed and calculated in previous works.^{5,12} Time $t = 0$ is defined as the time corresponding to a distance of $+3.5 \text{ mm}$ between cork and bottle. This arbitrary time reference is used to reduce the temporal dispersion of data, which is particularly pronounced for the experimental ones. This definition of $t = 0$ allows the time reference to be set in the

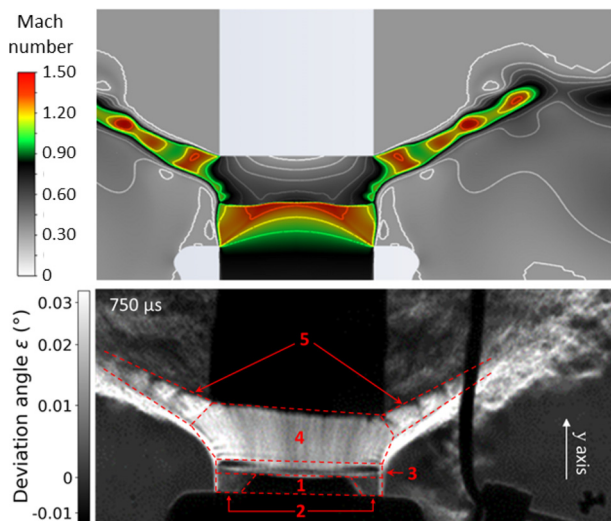


FIG. 6. Upper panel: Mach number map resulting from CFD simulation. Shades of gray are reserved for subsonic regions and the colors are reserved for supersonic regions. The solid green line represents the sonic line. Bottom panel: schlieren image at $750 \mu\text{s}$ after the gas has escaped from the bottle. Expansion zones discussed in the text are highlighted and labeled in red. Refractive index gradients are observed along the y axis.

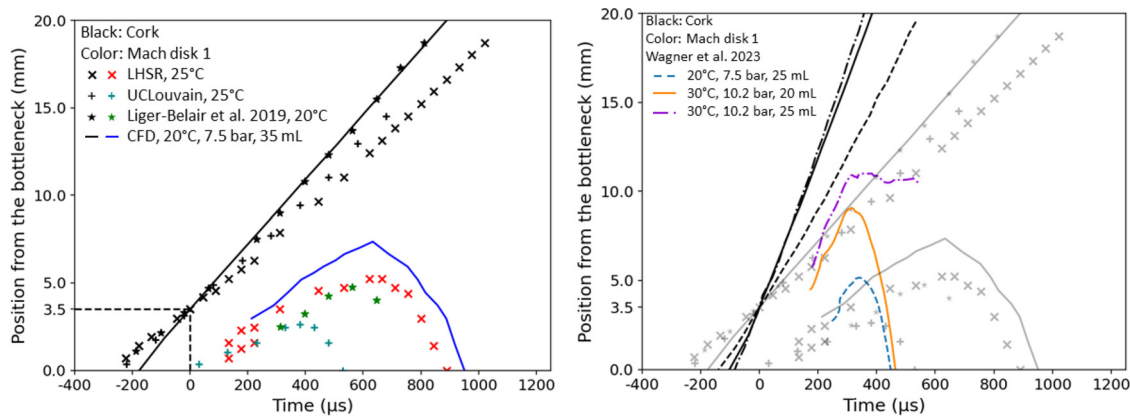


FIG. 7. Time evolution of the positions of the cork (in black) and the first Mach disk (in color). LHSR, UCLouvain, and CFD refer to videos related to Figs. 4, 5, and 3, respectively. Liger-Belair *et al.* data are reproduced from standard high-speed imaging.⁵ Wagner *et al.* are data reproduced from earlier numerical simulations.¹² The data presented in the left panel are reproduced in gray in the right panel. Initial thermodynamic conditions are indicated when available.

time domain where the cork motion is linear, so that the time reference no longer depends on the complex early dynamics of the cork. In this figure, the colored curves and symbols correspond to the Mach disk positions, while the black curves and symbols correspond to the positions of the cork. Symbols refer to experimental data, while solid lines refer to numerical simulations. For the sake of clarity, the existing data are divided in two different panels, but the data of the left panel are reproduced on the right panel in light gray for comparison purposes. When examining Fig. 7, readers should consider the frame rate of experimental data as an indication of their uncertainties over time ($\Delta t = 44, 50$, and $83 \mu\text{s}$ for LHSR, UCLouvain, and Liger-Belair *et al.*⁵ studies, respectively).

Figure 7 clearly shows that the cork reaches a constant speed around $100 \mu\text{s}$ after opening. This figure also highlights the experimental reproducibility of the time evolution of the first Mach disk, particularly before the Mach disk reaches its maximum distance value. Larger deviations are observed when focusing on the maximum height, which corresponds to the beginning of the last stage (expansion collapse stage). They probably result from variations in initial thermodynamic conditions (pressure, temperature, composition of the gas/champagne). No straightforward relationship between cork speed and maximum Mach disk height could be established: Compared with UCLouvain measurements, the cork speed at LHSR is slightly slower, while the maximum Mach disk height is greater. The CFD simulation presented in this work is also shown in Fig. 7 (continuous blue curve) and correctly reproduces the global evolution of the Mach disk. The recent full fluid–structure interaction model including the friction of the cork in the bottle¹² is also compared to our results in the right panel of Fig. 7. This model shows greater disagreement with our experimental observations. The systematic overestimation of the cork speed could be the reason for the faster time evolution of the Mach disk.

Finally, besides the supersonic dynamics described above, damped periodic gas breathing was observed with a frequency of ~ 600 to 1000 Hz for large times $> 2 \text{ ms}$ (see video UCLouvain, Multimedia view). This damped periodic oscillation at the bottleneck outlet has already been observed in previous works.^{4,5} This observation is in agreement with previous simulations^{10,12} where a compression wave

propagates in the bottle on this large timescale. According to these simulations, Mach disks 1 and 2 convert into a compression wave that propagates back in the bottle. This compression wave is reflected by the liquid surface and at the bottleneck outlet (due to acoustic impedance mismatch). Gas dynamics at the bottleneck outlet over a long period was simulated by Wagner *et al.*¹² and an oscillation period of $\sim 1 \text{ ms}$ was calculated. This oscillation period concurs with our observation.

IV. CONCLUSIONS AND PERSPECTIVES

In this study, an extensive characterization of the rich fluid dynamics following the uncorking of a bottle of champagne was carried out. Based on CFD calculation relying on the overset meshing method, two types of supersonic expansions were predicted: an impinging supersonic expansion along the symmetry axis of the bottle and a free radial supersonic expansion on side of the cork. High-speed schlieren imaging has been used to test the qualitative validity of these predictions. The presence of diamond shock cells in the radial gas expansion zone was observed experimentally for the first time, confirming the supersonic nature of the deflected jet. The supersonic nature of the axial expansion is also confirmed by identifying the supersonic and subsonic zones of this expansion, in agreement with the CFD calculations. Beyond the confirmation of the two supersonic expansions, formation of a second Mach disk during the collapse of the supersonic expansion has been revealed by schlieren imaging. This second Mach disk was predicted in our new CFD simulations and was predicted in a recent work¹² but was never observed experimentally. In addition, horizontal ripples in the outer part of the deflected jet were observed but not predicted in any previous CFD calculation. In parallel to this work, some of us have evidenced aggregation from an expansion of a mixture of CO_2 and H_2O .^{5,35} Evidencing such molecular mechanisms in real time would provide a fresh perspective on this interesting process.

SUPPLEMENTARY MATERIAL

See the supplementary material for additional results and explanations: Fig. S1 for illustration of the flow turbulence; Fig. S2 for aging

of champagne corks; Fig. S3 for calibration of the deviation angles in schlieren imaging; Fig. S4 for ray tracing for a high-density zone in an axially symmetric expansion; Fig. S5 for illustration of Fig. 3 with a common density scale; Fig. S6 for presentation of a reference cork launcher; Fig. S7 for schlieren images using the reference cork launcher; and Fig. S8 for qualitative comparison between observed and simulated ripples.

ACKNOWLEDGMENTS

The authors acknowledge the financial support provided by the Fédération Wallonie Bruxelles, through the project “iBEAM” funded by its research program “Actions de Recherche Concertées” (Convention No. 18/23-090) and the FNRS for a “mandat d’impulsion scientifique” (MIS, Grant No. F.4038.24). B.T. is a research fellow (ASP) of the FNRS. The authors also acknowledge the Institute of Mechanics, Materials, and Civil Engineering of UCLouvain for its technical support. The Castelnau Champagne producer is thanked for providing bottles used in one of the measurement campaigns.

AUTHOR DECLARATIONS

Conflict of Interest

The authors have no conflicts to disclose.

Author Contributions

J. N. Fréreau and B. Tomasetti contributed equally to this work.

J. N. Fréreau: Conceptualization (equal); Data curation (equal); Investigation (equal); Validation (equal); Writing – original draft (equal); Writing – review & editing (equal). **B. Tomasetti:** Conceptualization (equal); Data curation (equal); Investigation (equal); Validation (equal); Writing – original draft (lead); Writing – review & editing (lead). **A. Benidar:** Conceptualization (equal); Data curation (equal); Writing – review & editing (equal). **G. Liger-Belair:** Conceptualization (equal); Data curation (equal); Investigation (equal); Writing – original draft (equal); Writing – review & editing (equal). **A. B. Himakar:** Data curation (equal). **M. P. Lakhani:** Data curation (equal). **R. P. Srikanth:** Data curation (equal). **S. V. Nikhilesh:** Validation (equal). **J. Gopalan:** Validation (equal). **S. M. V. Rao:** Validation (equal). **R. Georges:** Supervision (equal). **C. Lauzin:** Supervision (equal); Validation (equal).

DATA AVAILABILITY

The data that support the findings of this study are available within the article and its supplementary material.

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